#### PH253 Lecture 4: relativity Lorentz Transformations, Dynamics

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### Outline



#### 2 Doppler shift







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#### Lorentz transformations between coordinate systems

$$x' = \gamma (x - \nu t)$$
 or  $x = \gamma (x' + \nu t')$  (1)

$$\mathbf{y}' = \mathbf{y} \tag{2}$$

$$z' = z \tag{3}$$

$$t' = \gamma \left( t - \frac{\nu x}{c^2} \right)$$
 or  $t = \gamma \left( t' + \frac{\nu x'}{c^2} \right)$  (4)

Here (x, y, z, t) is the position and time of an event as measured by an observer in O stationary to it. A second observer in O', moving at velocity v along the x axis, measures the same event to be at position and time (x', y', z', t').



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# Velocity Addition

We have an observer in a frame O, and a second observer in another frame O' who are moving relative to each other at a velocity v. Both observers measure the velocity of another object in their own frames  $(v_{obj} \text{ and } v'_{obj})$ . We can relate the velocities measured in the different frames as follows:

$$v_{obj} = \frac{v + v'_{obj}}{1 + \frac{v v'_{obj}}{c^2}} \qquad v'_{obj} = \frac{v_{obj} - v}{1 - \frac{v v_{obj}}{c^2}}$$
(5)

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Again,  $v_{obj}$  is the object's velocity as measured from the O reference frame, and  $v_{obj}$  is its velocity as measured from the O' reference frame.

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# Velocity Addition

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Again,  $v_{obj}$  is the object's velocity as measured from the O reference frame, and  $v_{obj}$  is its velocity as measured from the O' reference frame.

Decide if you should add or subtract velocities, then pick the formula.

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• What about velocities along y and *z*?



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$$\frac{\mathrm{d}t'}{\mathrm{d}t} = \frac{\mathrm{d}}{\mathrm{d}t} \left( \gamma \left( t - \frac{\nu x}{c^2} \right) \right) = \gamma \left( 1 - \frac{\nu \nu_x}{c^2} \right)$$

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(8)

Put it together:

$$\nu_{y}' = \frac{\nu_{y}}{\gamma \left(1 - \frac{\nu \nu_{x}}{c^{2}}\right)} \tag{9}$$



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Put it together:

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Transverse velocities do change! Similar expression along z.

This is a homework problem.



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### Outline



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• Classically: sound frequency higher for approaching source, lower for receding.



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- With wave speed *v*, source speed *v*<sub>s</sub>, and receiver speed *v*<sub>r</sub>:



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$$f' = f_o \left( \frac{\nu + \nu_r}{\nu - \nu_s} \right) \tag{10}$$



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Expect this for light, but  $v_{\text{light}} >> v_{\text{sound}}$  ... harder to observe

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- In stationary frame (e.g., next to light source), period is  $T = \frac{\lambda}{c}$
- In frame approaching light source at  $\nu_o,\,T'=\gamma T$
- *And*, crests of waves are coming faster:  $v'_w = v_w + v_o$



• In source frame,  $\lambda = v_w T$ 



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Moving *toward* source at  $v_0$  shortens wavelength  $\lambda$  toward blue



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Moving *toward* source at  $v_0$  shortens wavelength  $\lambda$  toward blue

Similarly, moving *away* source at  $v_0$  lengthens wavelength  $\lambda$  toward red ( $v_0 < 0$ )

Red shift: ratio of shifted wavelength to what a stationary observer sees:

$$z = \frac{\lambda'}{\lambda} = \sqrt{\frac{c + \nu_o}{c - \nu_o}}$$
(12)

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•  $v_o > 0$  when approaching,  $v_o < 0$  when receding.



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Relativity principle: makes no difference if source or observer is moving.


# Outline

Lorentz Transformations

## Doppler shift







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#### • How to conserve p in all frames? Whose velocity, whose time?



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proper velocity = 
$$\eta = \left(\frac{\text{rate of change of posn. in observer's frame}}{\text{proper time measured by moving observer}}\right)$$



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$$\eta = \frac{dl}{dt_p} = \frac{dl/dt}{dt_p/dt} = \frac{\nu}{dt_p/dt}$$
(13)  
$$\frac{dt_p}{dt} = \frac{d}{dt} (t/\gamma) = \frac{1}{\gamma}$$
(14)  
$$\implies \eta = \frac{\nu}{\sqrt{1 - \nu^2/c^2}} = \gamma \nu$$
(15)



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(15)

$$\implies \eta = \frac{v}{\sqrt{1 - v^2/c^2}} = \gamma v \tag{15}$$

Long story short:  $p = m\eta = \gamma m\nu$ 



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Maintains p conservation for all observers, e.g., collisions.



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What about mass? Modern view: *invariant* - all agree on rest mass



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Lorentz Transformations

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The work-energy theorem relates the change in kinetic energy of a particle to the work done on it by an external force:  $\Delta K = W = \int F dx$ . Writing Newton's second law as F = dp/dt, show that  $W = \int v dp$  and integrate by parts to obtain the result

$$K = \frac{\mathrm{mc}^2}{\sqrt{1 - \mathrm{v}^2/\mathrm{c}^2}} - \mathrm{mc}^2$$



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We are concerned with finding the kinetic energy of a particle accelerated from rest, starting with zero kinetic energy, so we may simply write  $K = W = \int F dx$  and drop the  $\Delta$ . Using what we have, and noting v = dx/dt:

$$K = \int_{v=0}^{v=v} F \, dx = \int_{0}^{v} \frac{dp}{dt} \, dx = \int_{0}^{v} dp \, \frac{dx}{dt} = \int_{0}^{v} v \, dp \tag{16}$$



We can integrate by parts. Let f = dp and g' = dx/dt, and recall integration by parts gives  $\int fg' = fg - \int f'g$ .

$$K = \int_{0}^{\nu} \nu \, dp = p\nu \Big|_{0}^{\nu} - \int_{0}^{\nu} p \, d\nu = p\nu - \int_{0}^{\nu} p \, d\nu \tag{17}$$



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Noting  $p = mv / \sqrt{1 - v^2 / c^2}$ ,



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Noting  $p = mv / \sqrt{1 - v^2 / c^2}$ ,

$$K = \frac{mv^2}{\sqrt{1 - v^2/c^2}} - \int_0^v \frac{mv}{\sqrt{1 - v^2/c^2}} \, dv$$
(18)  
$$= \frac{mv^2}{\sqrt{1 - v^2/c^2}} + mc^2 \sqrt{1 - v^2/c^2} \Big|_0^v$$
(19)

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#### Continuing:

$$K = \frac{mv^2}{\sqrt{1 - v^2/c^2}} + mc^2\sqrt{1 - v^2/c^2} - mc^2$$
(20)

$$=\frac{mv^{2}+mc^{2}\left(1-v^{2}/c^{2}\right)}{\sqrt{1-v^{2}/c^{2}}}-mc^{2}$$
(21)

$$= \frac{\mathrm{mc}^2}{\sqrt{1 - \mathrm{v}^2/\mathrm{c}^2}} - \mathrm{mc}^2 = \left(\frac{1}{\sqrt{1 - \mathrm{v}^2/\mathrm{c}^2}} - 1\right) \mathrm{mc}^2 \qquad (22)$$

$$\implies \mathsf{K} = (\gamma - 1)\,\mathsf{mc}^2\tag{23}$$



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# Low v? For v << c, $\gamma = (1 - v^2/c^2)^{-1/2} \approx 1 + \frac{v^2}{c^2}$



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(Binomial -  $(1 + x)^n \approx 1 + nx$  for  $x \ll 1$ )



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(Binomial -  $(1 + x)^n \approx 1 + nx$  for  $x \ll 1$ )  
Plug this in:  $K = (\gamma - 1) mc^2 \approx (1 + \frac{v^2}{c^2} - 1) mc^2 = \frac{1}{2}mv^2$ 



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#### • Now if $K = (\gamma - 1) \operatorname{mc}^2$ , as $\nu \to c$ then $K \to \infty$ .



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- Only massless objects (light) can travel at c



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- Takes infinite energy to reach light speed, not happening
- Only massless objects (light) can travel at c
- What about second term, independent of speed?!?

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$$K = (\gamma - 1) \operatorname{mc}^2 = \gamma \operatorname{mc}^2 - \operatorname{mc}^2$$



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- $K = (\gamma 1) \text{ mc}^2 = \gamma \text{mc}^2 \text{mc}^2$
- Subtract off a constant energy? Implies  $K = E_{total} E_{rest}$



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- $K = (\gamma 1) \text{ mc}^2 = \gamma \text{mc}^2 \text{mc}^2$
- Subtract off a constant energy? Implies  $K = E_{total} E_{rest}$
- Or,  $E_{tot} = K + mc^2 = \gamma mc^2$  and  $E_{rest} = mc^2$



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- Bodies have *rest energy* associated with mass mass-energy equivalent!

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Can go further ...

$$E_{tot}^{2} - (pc)^{2} = (mc^{2})^{2}$$
(24)  
$$E_{tot}\nu = pc^{2}$$
(25)

Better form: relativistic E - p relationship



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 Image: Second seco

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For massless objects, this implies:

$$p = \frac{E}{c}$$
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If you combine it all, you come to an even wilder conclusion.

p



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3

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c is really a speed limit for any massive particle

