PHYSICS 102

DR. LECLAIR

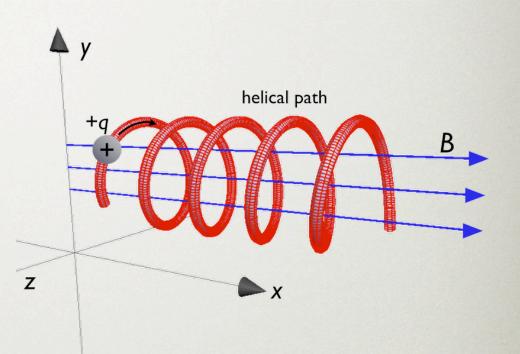
OFFICIAL THINGS

Lecture:

- 227 Gallalee
- every day! 10-11:45am

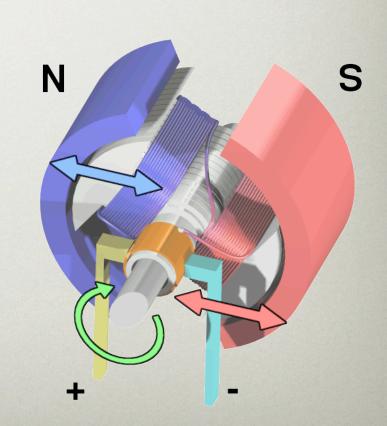
Lab:

- M-W-Th 3:30-6pm
- start 30mins later than scheduled
- usually done by 5-5:30, probably



OFFICIAL THINGS

- Dr. Patrick LeClair
 - pleclair@ua.edu please include 'ph102' in subject
 - office: 228 Bevill (348-0449)
 - lab: 180 Bevill
- Office hours:
 - 12-1pm in Gallalee
 - 1-2pm in Bevill
- other times by appointment



OFFICIAL THINGS

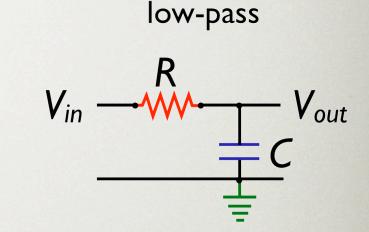
Graduate Assistants:

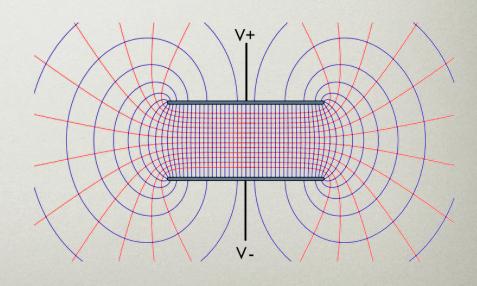
Tetyana Konak

Yinjun Zhang

Ru Zhu

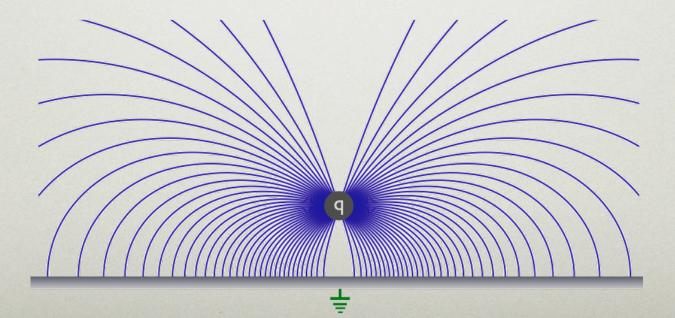
Physics Help Desk
 Hours TBA





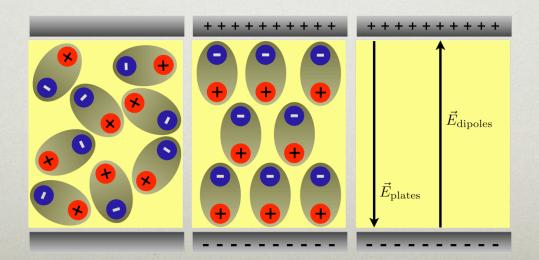
MISC. FORMAT ISSUES

- lecture and labs will try to stay linked
- learn a concept, then demonstrate it
- friday lecture: more problems + quiz
- working in groups is encouraged for homework



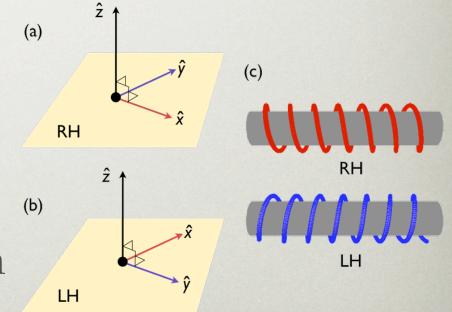
SOCIAL INTERACTION

- we need you in groups of 3-4 for labs
- groups are not assigned ...
 - ... so long as they remain functional relationships
 - even distribution of workload



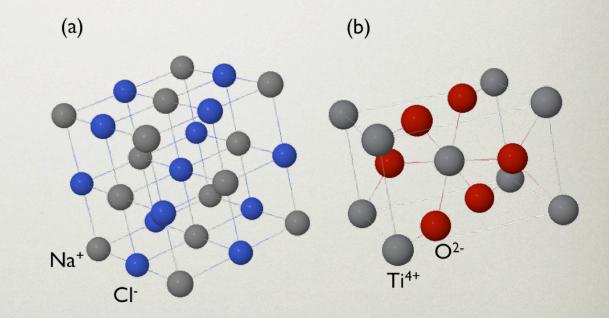
WHAT WILL WE COVER?

- relativity
- electric forces & fields
- electrical energy & capacitance
- current & resistance
- dc circuits
- magnetism
- electromagnetic induction
- ac circuits & EM waves



WHAT WILL WE COVER (CONT.)

- reflection and refraction
- mirrors & lenses
- wave optics
- quantum physics
- atomic physics
- nuclear physics



GRADING AND SO FORTH

- labs/exercises
- quizzes, homework
 - weekly
 - in-class questions
- exams
 - 1 multiple choice
 - 1 problem-based
 - 1 final

		%	
Component	Sections	section	total
In-class work	Labs & Exercises [†]	15	
	Quizzes	15	
			30
$Outside\ work$	Homework problems	15	
			15
Hour Exams	Exam I	15	
	Exam II	15	
			30
Final Exam			25

[‡] The lowest single grade will be dropped.

HOMEWORK

- out every monday on the blog [pdf]
- due the following monday at 3pm
- hard copy or email (e.g., scanned) both OK my Gallalee or Bevill mailbox give to TA at lab time
- can collaborate
- have to show your work to get credit.
 BUT turn in your own

QUIZZES

- every friday (at least)
- ~5-10 question multiple choice
- that week's work
- 10-15 min anticipated

occasionally and randomly in lecture

LABS / EXERCISES

- CRUCIAL: we will start closer to 3:15
- something due every day lab is held
- if not a "real" lab: in-class exercises or simulations
- drop 2 ...

USUALLY will not take 3 hours

STUFF YOU NEED

textbook

Serway & Faughn. get a used one.

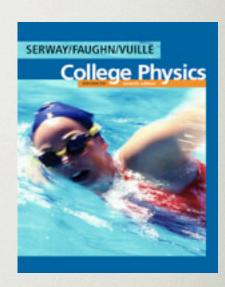
course notes (optional)

PDF online (do not print it here)

calculator

basic with trig/log

notebook



Current and Resistance

1 Electric Current

ELECTRIC current is something that we use and hear about every day, but few of us stop to thin
about what it really is. What is an electric current? An electric current is nothing more than the ne
flow of charges through some perion in a combuttor.

If we take a cross section of a conductor, such as a circular whee, an other current is said to circl of there is a neight free of chapes through this argine. It amount of current is simply the rate at which charge is flowing, the number charges per unit time that treverse the cross-sections. Strictly speaking, set to choose the cross-sections for defining charge flow such take the draps of per perpendicular to that surface, somewhat like we did for Gauss's law. Figure shows a curroon depiction of how we define current.

rows a currons repeated on one we tenther current.

Current is a fixe of charges through, a wire in the same way that water flow is fax of water through a pipe. As we shall see, this is a reasonable way to thisk on electric circuits as well—current adverse, but no flow somethers, and you n't want an open connection any more than your would want an open-cented terp pice. Voltage is mere like a pressure gauge — you can have a webrage even am nothing is flowing, it just means there is the potential for flow (northy punmental flow of the property of the contraction of the potential for flow (northy punmental flow of the potential for flow (northy punmental flow of the property of the potential for flow (northy punmental flow of the potential for flow (northy punmental flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential flow of the potential for flow (northy punmental flow of the potential fl



seams there is the potential for flow (nerdy pun tionship between current, so age, and resistance. ¹⁴

2 flows perpendicularly through a particular

Electric Current: if a net amount of charge ΔQ flows perpendicularly through a surface

interval M, the electric current I is:

 $I = \frac{\Delta U_f}{\Delta t}$

This represents a conservation law as well. Charge can neither be created or distroyed. If we have some steady stream of charge pouning into of a region of fixed volume, then the charge density inside would continually grow (tending toward infinity!) if there were not also some compensating flow of charges out of the volume. Putting it the other way around, if a steady stream of charges were forwing the fixed volume, the charge density would also become infinitely lawre if there were not some other source of charges to reclaive

the change in the total number of charges in a volume at any time has to equa through that volume, otherwise we would require spontaneous generation of charge

Units of electric current I: Coulombs per second [C/s] or Amperes [J

¹We have waved our hands a bit here, since we should talk about current density and charge density, but the essential are the same

73

SHOWING UP

no make-up of in-class work or homework
 "acceptable" + documented gets you a BYE

missing an exam is seriously bad.
 acceptable reason => makeup or weight final

• lowest 2 labs are dropped. I don't want to know.

DISTRACTIONS

- cell phones
 - keep it on a quiet mode.
 - take the call outside if it is urgent

"no food/drink"

at least one break during each lecture

OTHER

Academic misconduct

- do your own work on quizzes & exams
- suspected violations referred to A & S
- teamwork encouraged on labs/homework

Accessibility/disability accommodations

- for a request 348-4285 Disabilities services
- after initial arrangements, contact me

INTERNETS

- we have our own intertubes:
 - http://ph102.blogspot.com
 - updated several times a week. often late
 - comments (anonymous even) allowed
 - rss feed
- google calendar
- Facebook group ...
 - can add RSS feed of blog to facebook
- check blog & calendar before class

LET'S GET AT IT

The pace will have to be brutal.

Today & tomorrow

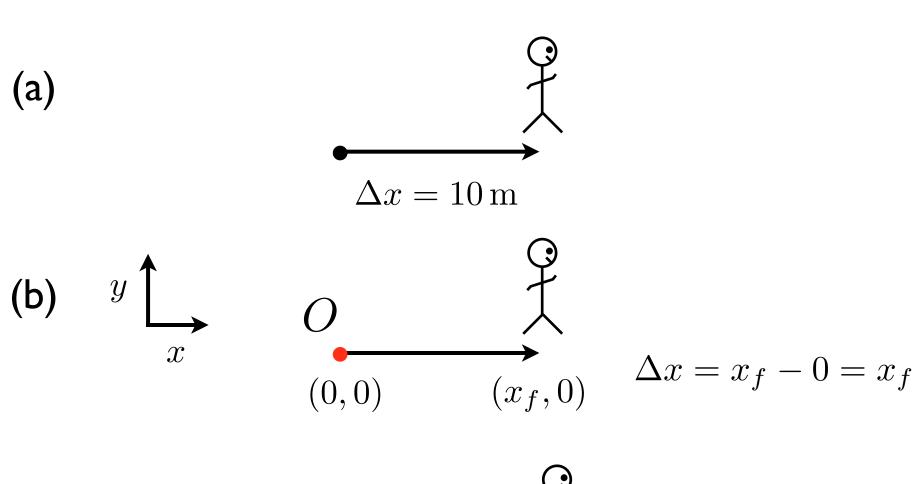
• Relativity (S & F ch. 26, Notes Ch. 2)

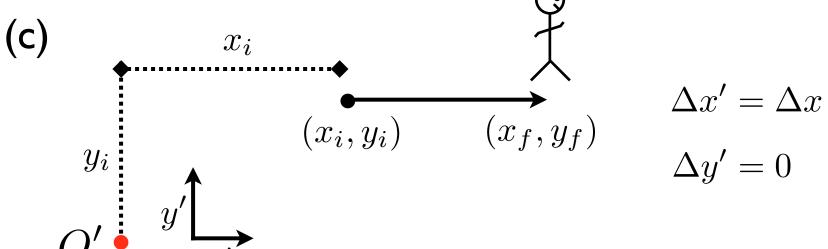
Thursday & Friday:

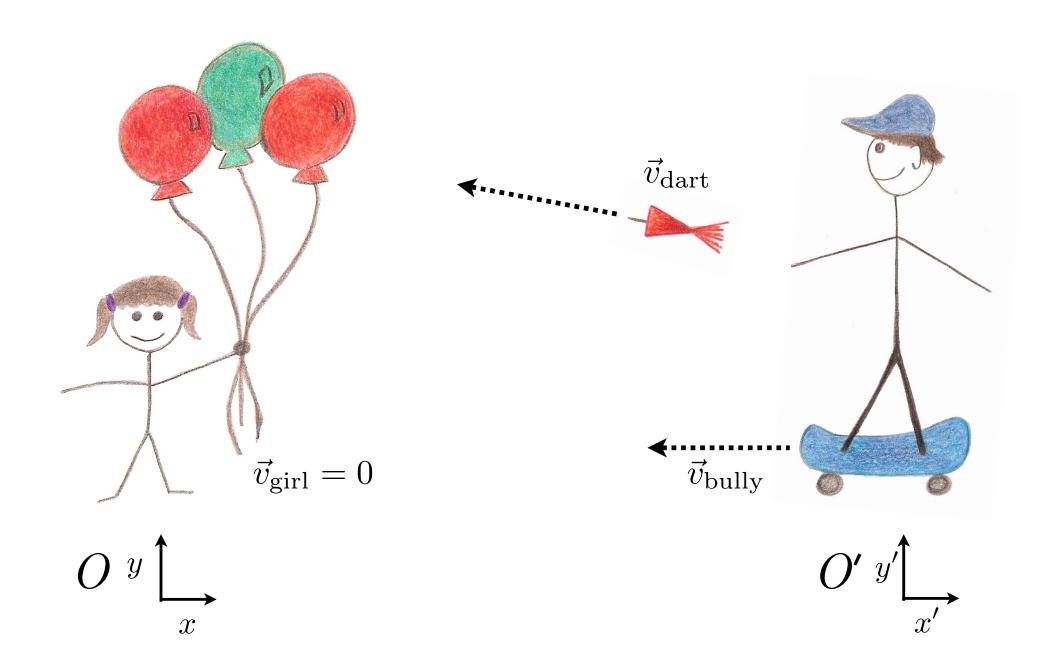
electric fields & forces

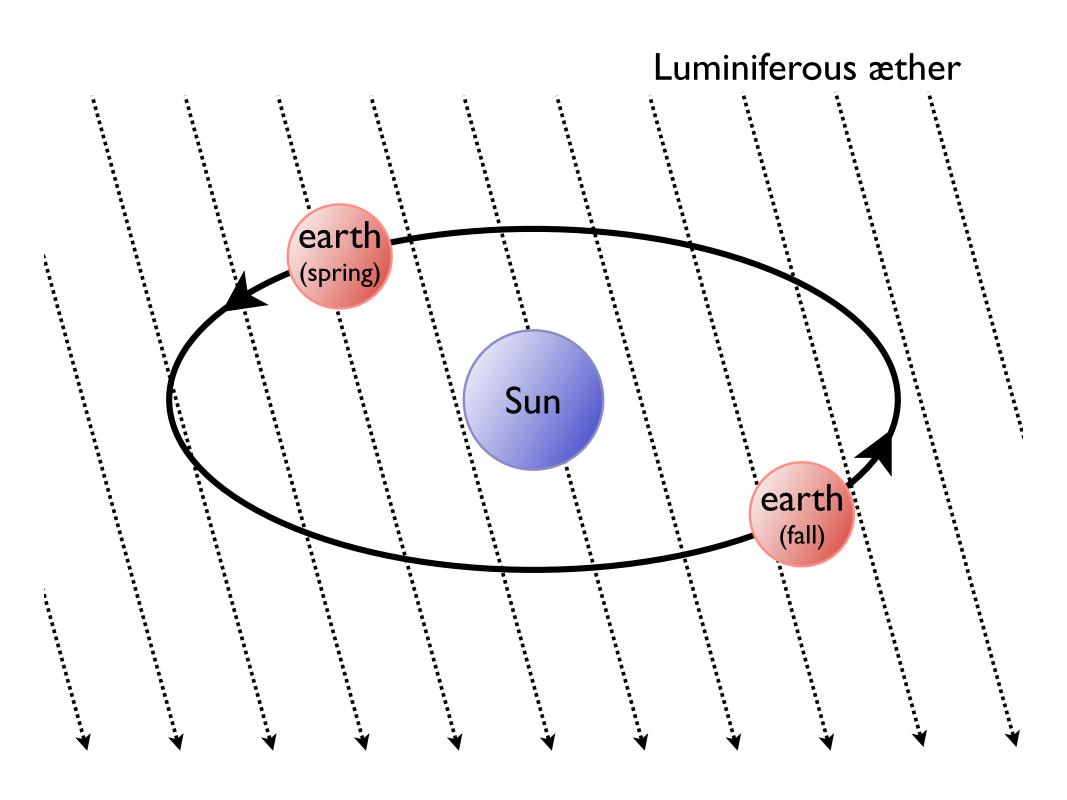
Tomorrow's lab: 3:30pm!!

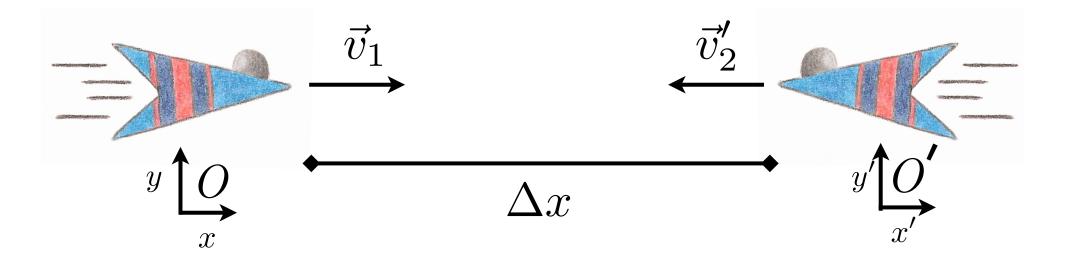
• research & writing assignment (yes really)

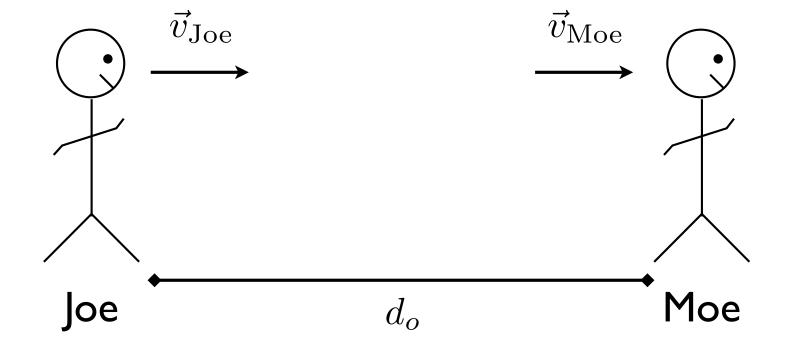






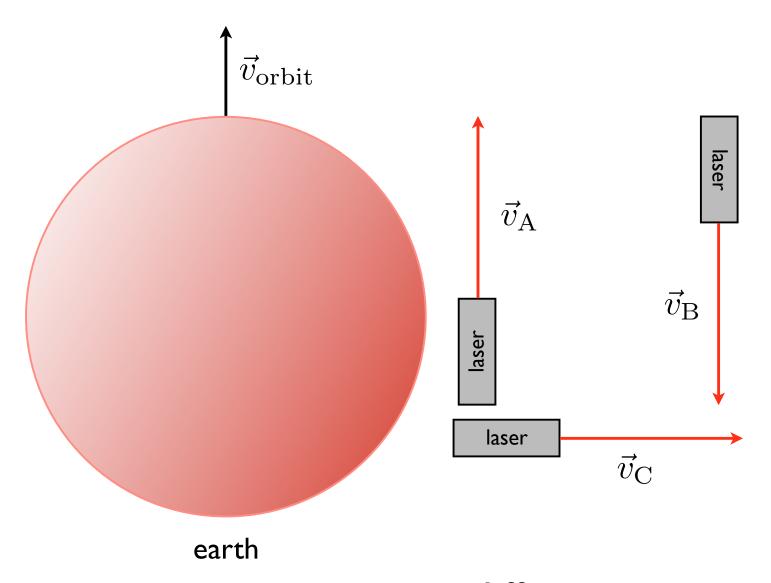




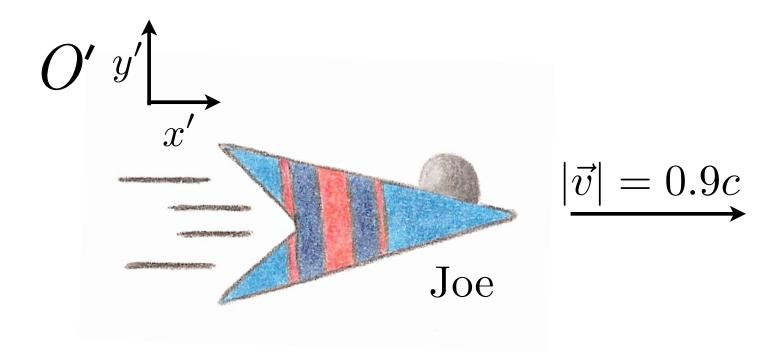


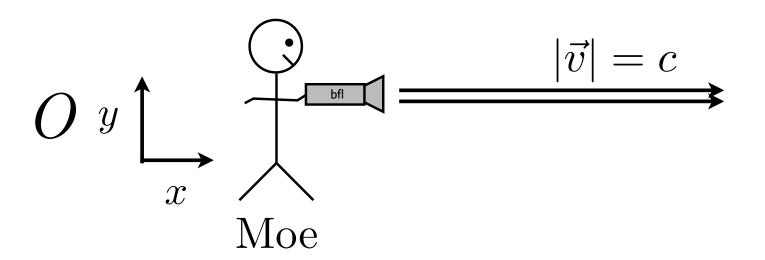
Choosing a coordinate system:

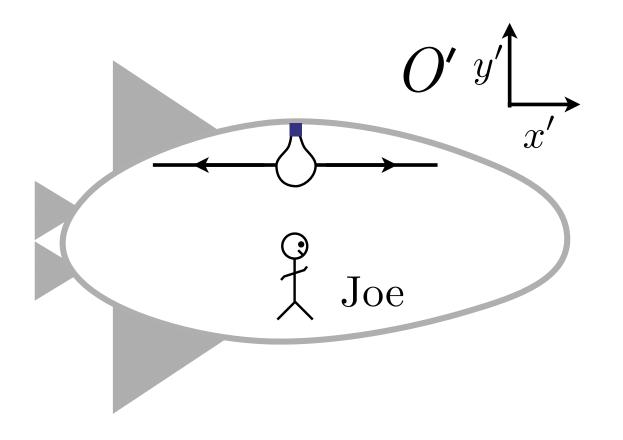
- 1. Choose an origin. This may coincide with a special point or object given in the problem for instance, right at an observer's position, or halfway between two observers. Make it convenient!
- 2. Choose a set of axes, such as rectangular or polar. The simplest are usually rectangular or Cartesian x-y-z, though your choice should fit the symmetry of the problem given if your problem has circular symmetry, rectangular coordinates may make life difficult.
- 3. Align the axes. Again, make it convenient for instance, align your x axis along a line connecting two special points in the problem. Sometimes a thoughtful but less obvious choice may save you a lot of math!
- 4. Choose which directions are positive and negative. This choice is arbitrary, in the end, so choose the least confusing convention.

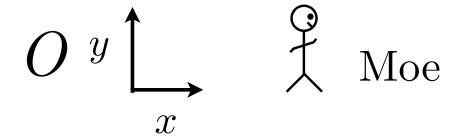


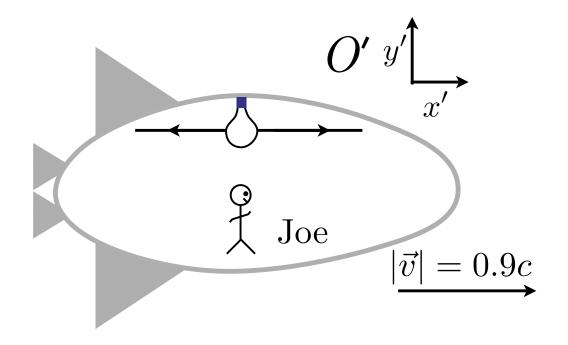
no difference can't measure earth's velocity relative to empty space

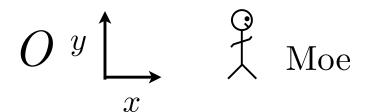




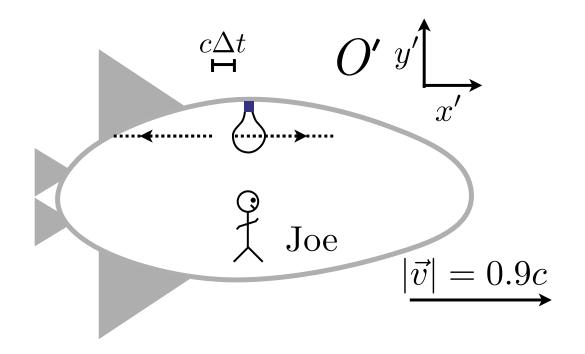








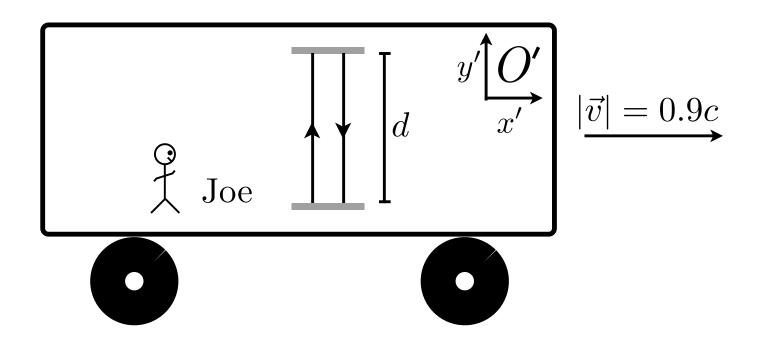
Joe flips on the light he sees the light hit the walls at the same time



$$O^y \xrightarrow{x} \stackrel{\circ}{\downarrow} Moe$$

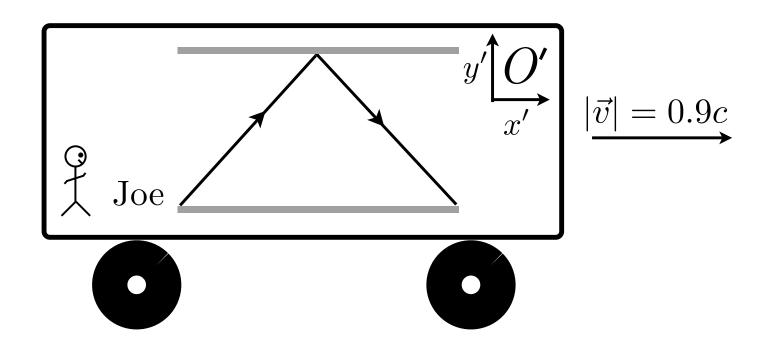
What does Moe see?

the ship moved
the origin of the light did not

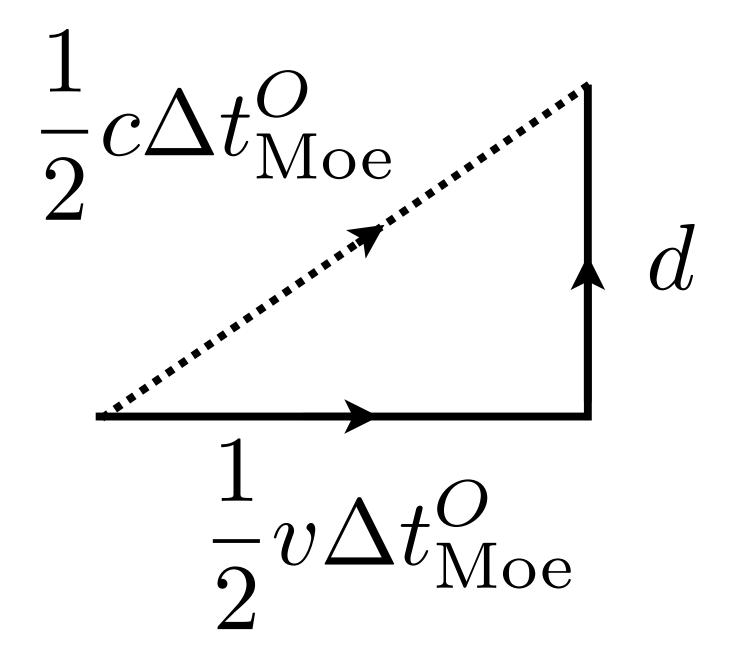


$$y
ightharpoonup O$$
 Moe

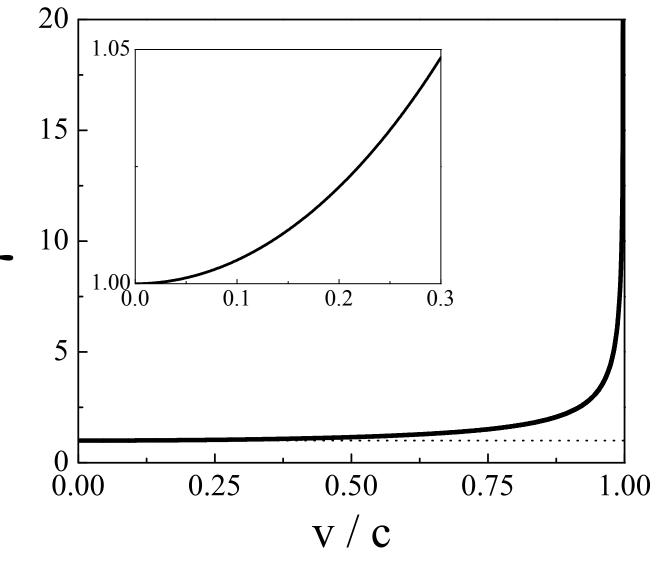
Joe bounces a laser off of some mirrors he counts the round trips this measures distance

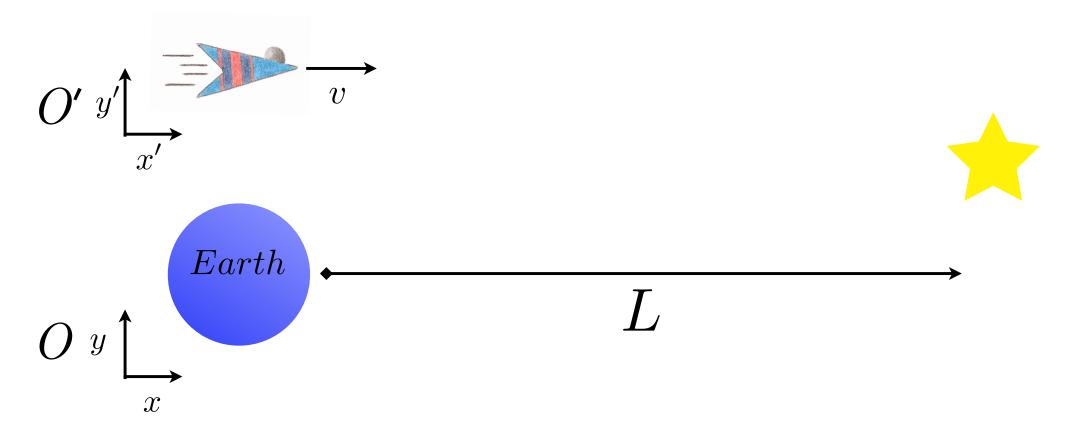


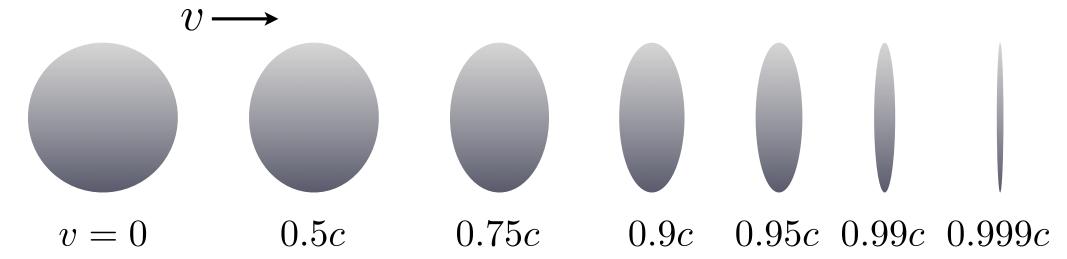
Moe sees the boxcar move once the light is created, it does not Moe sees a triangle wave

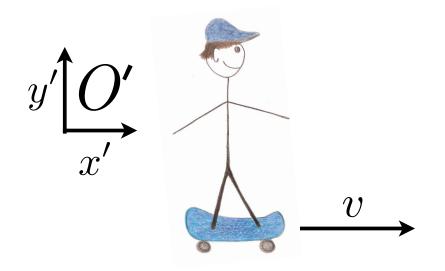


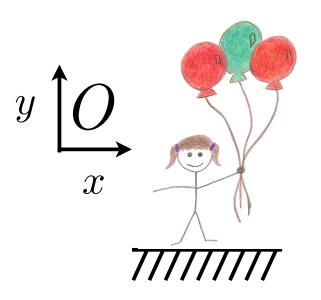
v [m/s]	$\frac{v}{c}$	γ	$1/\gamma$
0	0	0	∞
3×10^{6}	0.01	1.00005	0.99995
3×10^{7}	0.1	1.005	0.995
6×10^{7}	0.2	1.02	0.980
1.5×10^{8}	0.5	1.16	0.866
2.25×10^{8}	0.75	1.51	0.661
2.7×10^{8}	0.9	2.29	0.436
2.85×10^{8}	0.95	3.20	0.312
2.97×10^{8}	0.99	7.09	0.141
2.983×10^{8}	0.995	10.0	0.0999
2.995×10^{8}	0.999	22.4	0.0447
2.996×10^{8}	0.9995	31.6	0.0316
2.998×10^{8}	0.9999	70.7	0.0141
c	1	∞	0





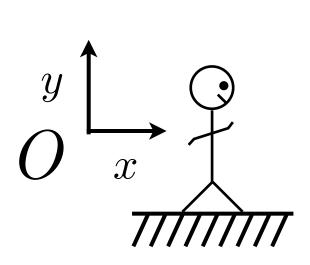


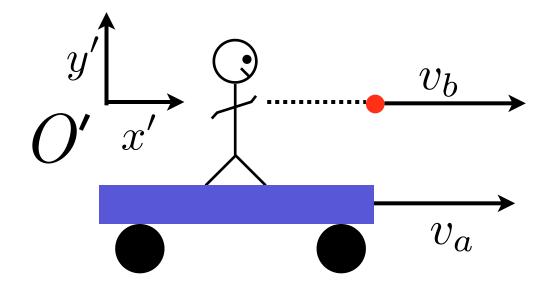


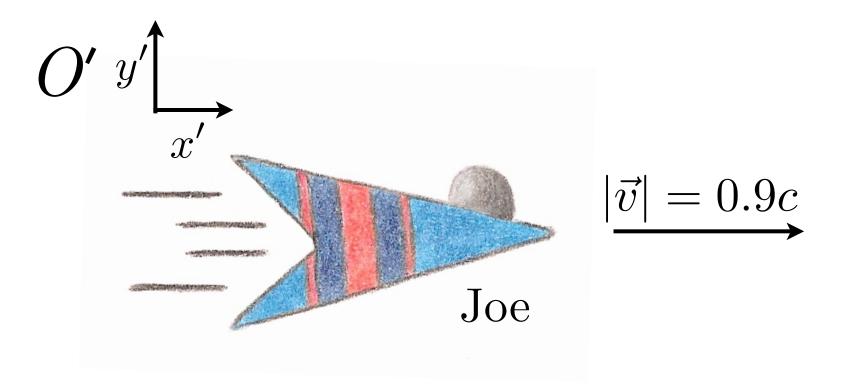


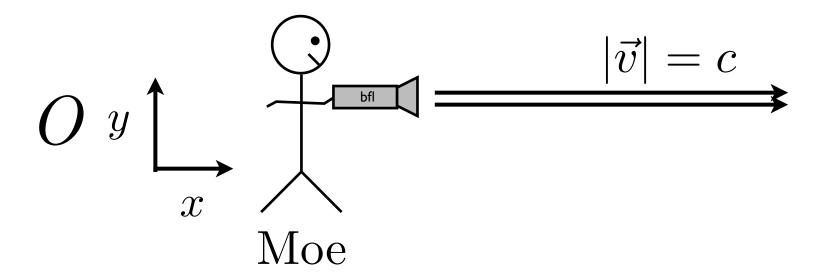


 ${\mathcal X}$









let's work out some problems

	foot 3.0 times per second. What is the frequency Hint: be careful about the difference between time
○ 5.0 taps/sec	
○ 6.7 taps/sec	
○ 1.8 taps/sec	
○ 3.0 taps/sec	
	speed. How do experimenters on earth measure those in the spaceship measure a clock on earth
○ slow; fast	
O slow; slow	
○ fast; slow	
○ fast; fast	
difference in your pulse rate? In the pulse rate O no; yes	speed relative to the earth, would you notice a of the people back on earth?
O no; no	
○ yes; no	
○ yes; no○ yes; yes	
yes; yesThe period of a pendulum is measured to be	be 3.00 in its own reference frame. What is the a speed of 0.950c with respect to the pendulum?
yes; yesThe period of a pendulum is measured to be	
 yes; yes The period of a pendulum is measured to be period as measured by an observer moving at a 	
 yes; yes The period of a pendulum is measured to be period as measured by an observer moving at a 6.00 sec 	

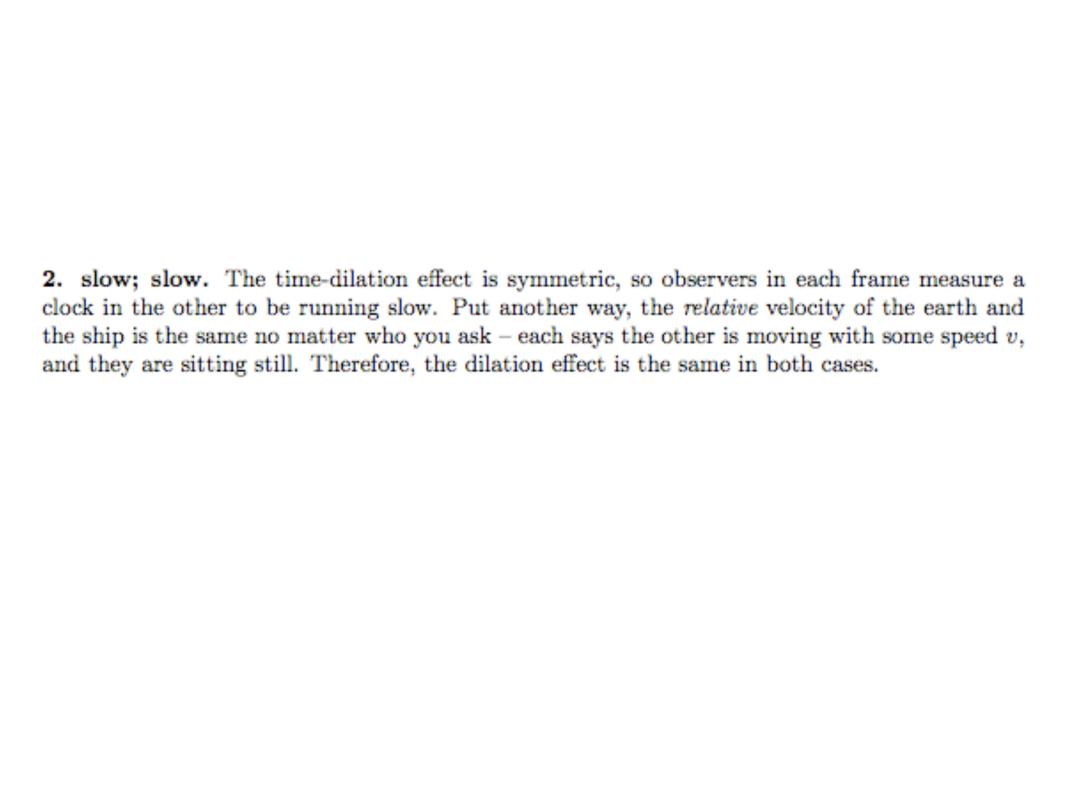
1.8 taps/sec. The 'proper time' Δt_p is that measured by the astronaut herself, which is 1/3 of a second between taps (so that there are 3 taps per second). The time interval between taps measured on earth is dilated (longer), so there are less taps per second. For the astronaut:

$$\Delta t_p = rac{1 \, \mathrm{s}}{3 \, \mathrm{taps}}$$

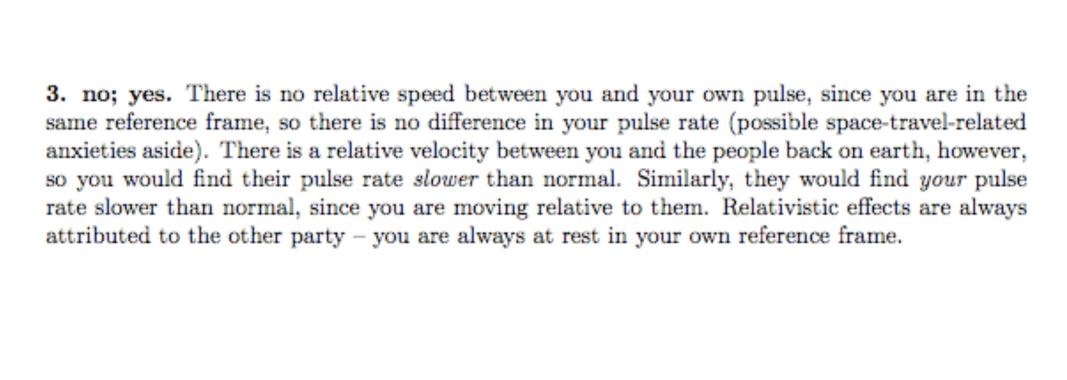
On earth, we measure the dilated time:

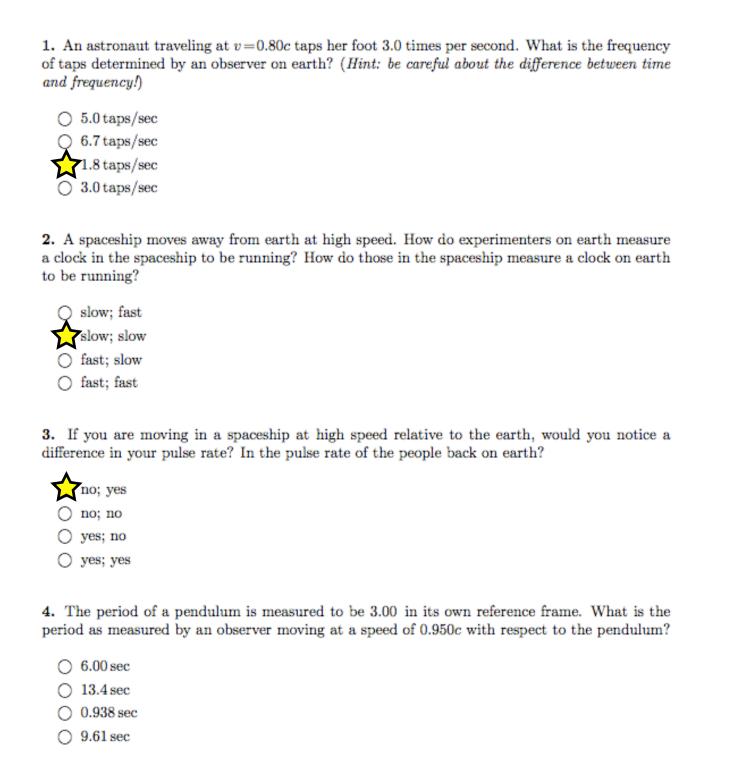
$$\Delta t' = \gamma \Delta t_p = \frac{1}{\sqrt{1 - \frac{0.8^2c^2}{c^2}}} \cdot \left(\frac{1\,\mathrm{s}}{3\,\mathrm{taps}}\right) = \frac{1}{\sqrt{1 - 0.8^2}} \cdot \left(\frac{1\,\mathrm{s}}{3\,\mathrm{taps}}\right) \approx \frac{0.56\,\mathrm{s}}{\mathrm{tap}} = \frac{1\,\mathrm{s}}{1.8\,\mathrm{taps}}$$

1. An astronaut traveling at $v = 0.80c$ taps her foot 3.0 times per second. What is the frequency of taps determined by an observer on earth? (Hint: be careful about the difference between time and frequency!)	
○ 5.0 taps/sec	
O 6.7 taps/sec	
1.8 taps/sec	
○ 3.0 taps/sec	
2. A spaceship moves away from earth at high speed. How do experimenters on earth measure a clock in the spaceship to be running? How do those in the spaceship measure a clock on earth to be running?	
○ slow; fast	
○ slow; slow	
○ fast; slow	
O fast; fast	
 3. If you are moving in a spaceship at high speed relative to the earth, would you notice a difference in your pulse rate? In the pulse rate of the people back on earth? \(\cap \) no; yes \(\cap \) no; no \(\cap \) yes; no \(\cap \) yes; yes 	
4. The period of a pendulum is measured to be 3.00 in its own reference frame. What is the period as measured by an observer moving at a speed of 0.950c with respect to the pendulum?	
○ 6.00 sec	
○ 13.4 sec	
○ 0.938 sec	
○ 9.61 sec	



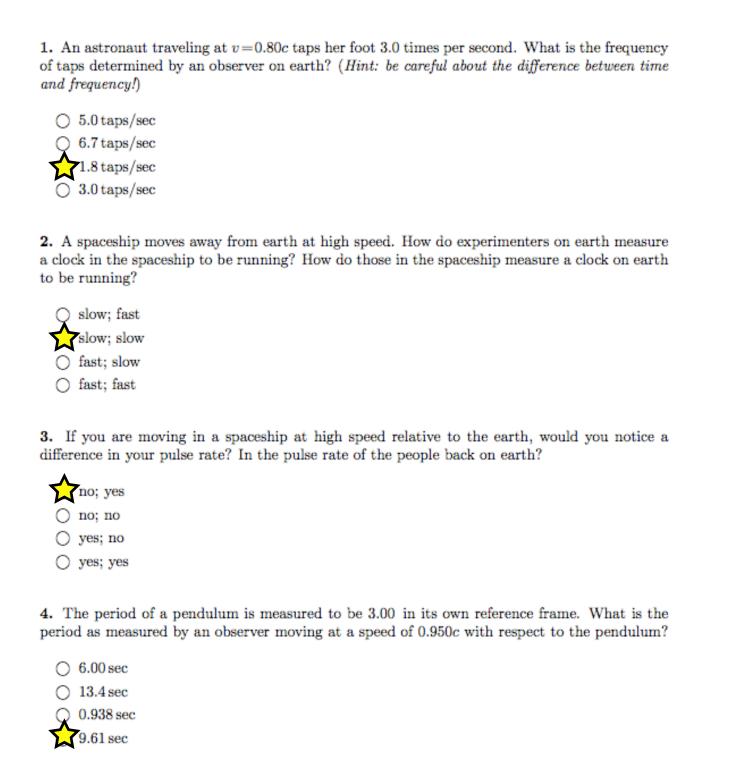
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○ 5.0 taps/sec
1.8 taps/sec
○ 6.7 taps/sec 1.8 taps/sec ○ 3.0 taps/sec
2. A spaceship moves away from earth at high speed. How do experimenters on earth measure a clock in the spaceship to be running? How do those in the spaceship measure a clock on earth to be running?
○ slow; fast
slow; fast slow; slow
O fast; slow
O fast; fast
3. If you are moving in a spaceship at high speed relative to the earth, would you notice a difference in your pulse rate? In the pulse rate of the people back on earth?
ono; yes
O no; no
○ yes; no
○ yes; yes
O yes; yes
O yes; yes 4. The period of a pendulum is measured to be 3.00 in its own reference frame. What is the period as measured by an observer moving at a speed of 0.950c with respect to the pendulum?
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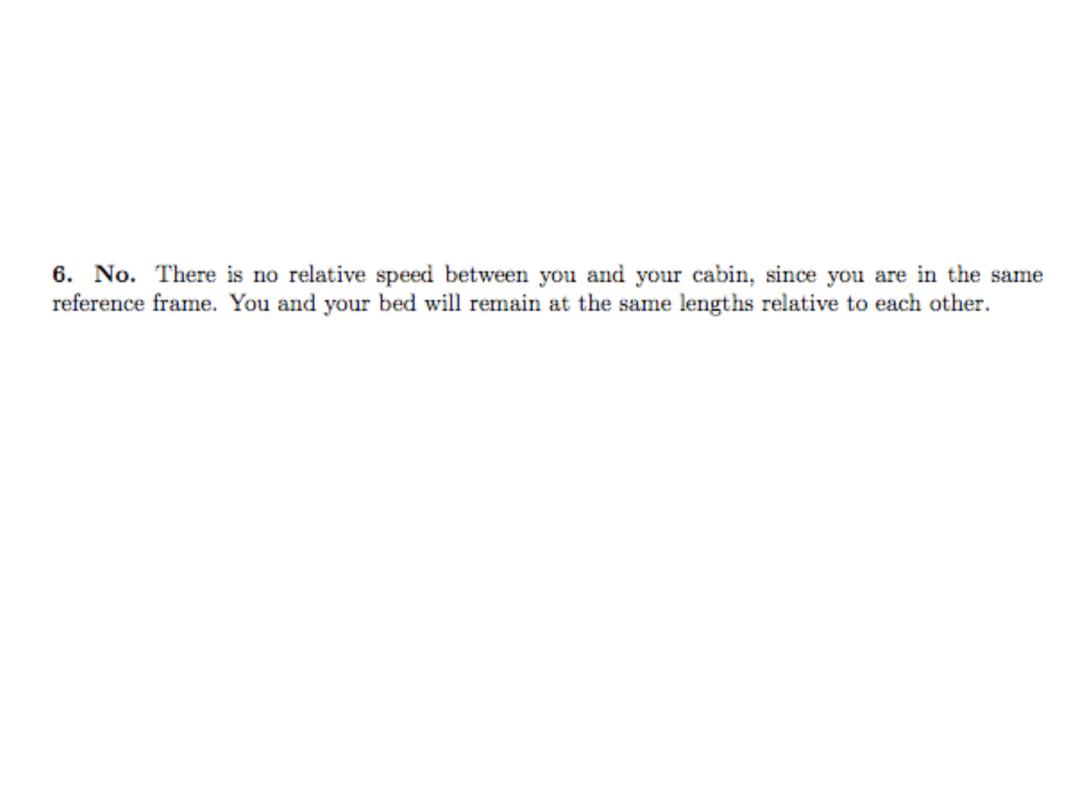


4. 9.61 sec. The proper time is that measured by in the reference frame of the pendulum itself, Δt_p = 3.00 sec. The moving observer has to observe a longer period for the pendulum, since from the observer's point of view, the pendulum is moving relative to it. Observers always perceive clocks moving relative to them as running slow. The factor between the two times is just γ:

$$\Delta t' = \gamma \Delta t_p = \frac{3.0 \sec}{\sqrt{1 - \frac{0.95^2 c^2}{c^2}}} = \frac{3.0 \sec}{\sqrt{1 - 0.95^2}} \approx 9.61 \sec$$



- 6. You are packing for a trip to another star, and on your journey you will travel at 0.99c. Can you sleep in a smaller cabin than usual, because you will be shorter when you lie down? Explain your answer.
- 7. A deep-space probe moves away from Earth with a speed of 0.88c. An antenna on the probe requires 4.0 s, in probe time, to rotate through 1.0 rev. How much time is required for 1.0 rev according to an observer on Earth?
- 8. A friend in a spaceship travels past you at a high speed. He tells you that his ship is 24 m long and that the identical ship you are sitting in is 18 m long.
- (a) According to your observations, how long is your ship?
- (b) According to your observations, how long is his ship?
- (c) According to your observations, what is the speed of your friend's ship?



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7. 8.42 s. The time interval in the probe's reference frame is the proper one Δt_p ... which makes sense, since the antenna is part of the probe itself! The probe and antenna are moving relative to the earth, and therefore the earthbound observer measures a longer, dilated time interval Δt':

probe =
$$\Delta t_p$$

earth = $\Delta t'$
 $\Delta t'$ = $\gamma \Delta t_p$

As usual, we first need to calculate γ . No problem, given the probe's velocity of 0.88c relative to earth:

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{1}{\sqrt{1 - \frac{(0.88c)^2}{c^2}}} = \frac{1}{\sqrt{1 - 0.88^2}} = 2.11$$

The proper time interval for one revolution Δt_p in the probe's reference frame is 4.0 s, so we can readily calculate the time interval observed by the earthbound observer:

$$\Delta t' = \gamma \Delta t_p = 2.11 \cdot (4.0 \, s) = 8.42 \, s$$

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8. 24 m; 18 m; 0.661c. Once again: if you are observing something in your own reference frame, there is no length contraction or time dilation. You always observe your own ship to be the same length. If your friend's ship is 24 m long, and yours is identical, you will measure it to be 24 m.

On the other hand, you are moving relative to his ship, so you would observe his ship to be length contracted, and measure a shorter length. Your friend, on the other hand, will observe exactly the same thing - he will see your ship contracted, by precisely the same amount. Your observation of his ship has to be the same as his observation of his ship - since you are only the two observers, and you both have the same relative velocity, you must observe the same length contraction. If he sees your ship as 18 m long, then you would also see his (identical) ship as 18 m long.

Given the relationship between the contracted and proper length, we can find the relative velocity easily. Your measurement of your own ship is the proper length L_p , while your measurement of your friend's ship is the contracted length L':

$$\begin{array}{rcl} L_p & = & \gamma L' \\ \Longrightarrow \gamma & = & \frac{L_p}{L'} = \frac{24}{18} = \frac{4}{3} \\ \\ \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} & = & \frac{4}{3} \\ \\ 1 - \frac{v^2}{c^2} & = & \frac{3^2}{4^2} = \frac{9}{16} \\ \\ \frac{v^2}{c^2} & = & 1 - \frac{9}{16} = \frac{7}{16} \\ \\ v & = & \sqrt{\frac{7}{16}}c = \frac{\sqrt{7}}{4}c \approx 0.661c \end{array}$$